

Deterministic Methods for the Boltzmann Equation

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AM281.1 Convection-Dominated Problems
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Outline

- 1 Introduction
 - The Equation
 - Collisions
- 2 Time-Step Splitting
 - Operator splitting
 - Runge-Kutta
- 3 Transport
 - Flux-Balance Methods
 - Semi-Lagrangian Methods
- 4 Collision Operator
 - Discretization
 - Bobylev-Rjasanow's Integral Transform Method
 - Pareschi-Russo's Spectral Method
 - Mouhot-Pareschi's Sub- N^6 Method

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The Boltzmann Equation

The Boltzmann Equation:

- Is an equation of *Statistical* Mechanics.
- Describes a rarefied gas.

Rarefied gas?

→ Enough space that particles in the same volume element may have differing velocities.

Contrast: *Continuum* Mechanics

Density Functions

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Density function:

$$\int_{\mathbb{R}^3 \times \mathbb{R}^3} f(v, x, t) dx dv = n(t).$$

The Boltzmann Equation

The Boltzmann equation is an Integro-PDE.

$$\frac{\partial f}{\partial t} + v \cdot \nabla_x f = \frac{1}{k_n} Q(f, f), \quad x, v \in \mathbb{R}^3$$

Also with Forces:

$$\frac{\partial f}{\partial t} + v \cdot \nabla_x f + F \cdot \nabla_v f = \frac{1}{k_n} Q(f, f), \quad x, v \in \mathbb{R}^3$$

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- LHS: particle transport according to v (“*Vlasov Equation*”)
- RHS: describes binary collisions between gas particles

Forces may come from: Maxwell’s Equations (Lorentz force),
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What is k_n ? → The *Knudsen Number*.

$$k_n := \frac{\lambda}{L} := \frac{\text{mean free path}}{\text{representative physical length scale}}$$

$k_n \rightarrow 0$:

Collision Term

The collision term splits into *gain* and *loss*.

$$Q(f, f) = Q^+(f, f) - L[f]f$$

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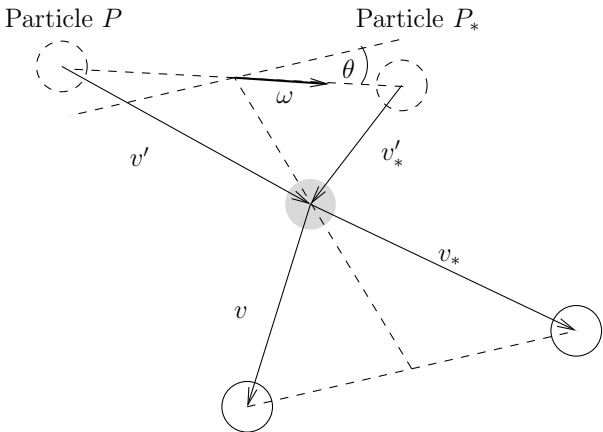
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Loss Term:

$$L[f] = \int_{\mathbb{R}^3} \int_{S^2} B(|v - v_*|, \cos \theta) f(v_*) d\omega dv_*.$$

Arggh. Too much notation. What are we describing, anyway?

Binary Collisions



Collision Term

Ok, let's look at this again.

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Loss Term:

$$L[f]f = \int_{\mathbb{R}^3} \int_{S^2} B(|v - v_*|, \cos \theta) f(v) f(v_*) d\omega dv_*$$

Q is a spatially local operator!

Collision Kernels

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Collision Kernels

- *Maxwellian gas*: $B(|v - v_*|, \cos \theta) = \text{const.}$
- *Hard Sphere gas*: $B(|v - v_*|, \cos \theta) = \text{const} \cdot |v - v_*|.$
- *Variable Hard Sphere (VHS) gas*:
 $B(|v - v_*|, \cos \theta) = \text{const} \cdot |v - v_*|^\alpha.$ (generalizes both cases above)

Of course, one can cook up many more complicated collision kernels.

Challenges

- Dimensionality: Eleven nested loops (naively) \rightarrow horrible!
- Timestepping: Different for transport and collision
 \rightarrow Splitting (next section)
- Shock Waves (Transport \rightarrow Section 3)

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- Dimensionality: Eleven nested loops (naively) \rightarrow horrible!
- Timestepping: Different for transport and collision \rightarrow Splitting (next section)
- Shock Waves (Transport \rightarrow Section 3)
- Collision operator: Deepest dimensionality. Can we do better than five nested loops at each point? (Collision \rightarrow Section 4)
- Positivity (if there is time)

Summary

Questions?

?

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General Framework

- Semi-discrete form of Boltzmann equation

$$\frac{dy}{dt} = f_1(y) + f_2(y)$$

- f_1 and f_2 may have significantly different properties
 - f_1 represents convection
 - f_2 is usually a stiff discretization of collision
- Possible nonlinearity of f_1 : high-order accurate explicit method
- Highly stiff f_2 : implicit method
- Conservation? Positivity?

Splitting Methods

- Strategy 1: operator splitting

- $S_1^{\Delta t}$ and $S_2^{\Delta t}$ are solution operators for

$$\frac{dy}{dt} = f_1(y) \qquad \frac{dy}{dt} = f_2(y), \qquad (1)$$

respectively

- Attempt to approximate $S^{\Delta t}$ with $S_1^{\Delta t}$ and $S_2^{\Delta t}$:

$$S^{\Delta t} \approx p(S_1^{\Delta t}, S_2^{\Delta t})$$

- Strategy 2: Runge-Kutta split methods

- Create two separate Runge-Kutta methods for each of the ODE's in (1), but evolve them together
- High-order accuracy requires coupling conditions between the two methods
- Disadvantage: requires semi-discrete form of PDE

Operator Splitting (1/2)

- First guess: sequential application

$$S^{\Delta t} \approx S_1^{\Delta t} S_2^{\Delta t}$$

- First-order in time
 - Very simple to formulate and elementary to implement
- “Strang splitting”

$$S^{\Delta t} \approx S_1^{\Delta t/2} S_2^{\Delta t} S_1^{\Delta t/2}$$

- Second-order in time
 - Not much more difficult than first-order splitting
 - Probably the most popular operator splitting algorithm used today

Operator Splitting (2/2)

- Higher-order accuracy?
 - Methods exist, but suffer from stability problems
 - Dia has developed 3rd, 4th order schemes

Runge-Kutta methods

- “Usual” s -stage Runge-Kutta methods have the form

$$y^{n+1} = y^n + \Delta t \sum_{j=1}^s b_j f(t^n + c_j \Delta t, Y_j)$$

$$Y_i = y^n + \Delta t \sum_{j=1}^s a_{ij} f(t^n + c_j \Delta t, Y_j)$$

where a_{ij} is an $s \times s$ matrix and b_j and c_j are $s \times 1$ matrices

- Can define explicit and implicit RK methods in this way

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Argh, More Splitting?

- Complexity: six-dimensional transport equation \rightarrow two 3-dimensional systems is more tractable

$$\frac{\partial f}{\partial t} = v \cdot \nabla_x f$$

$$\frac{\partial f}{\partial t} = F \cdot \nabla_v f$$

Flux-Balance Methods

- Finite-volume degrees of freedom:

$$f_i^n = \int_{I_i} f(x, t) dx$$

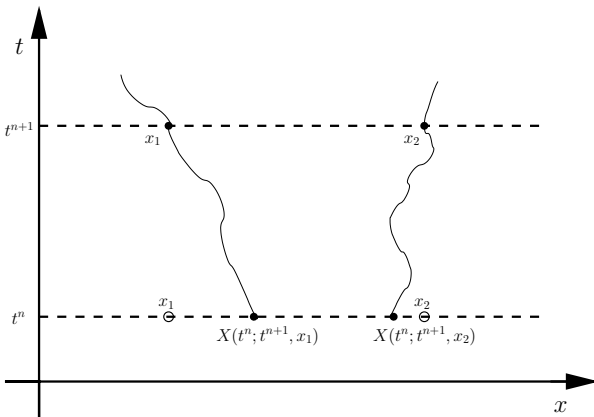
- Method of characteristics: f is constant along $x = X(t)$, where $X(t)$ solves the ODE

$$\frac{dX(s)}{ds} = v(s, X(s))$$

$$X(t^{n+1}) = x.$$

- $f(x, t^{n+1}) = f(X(t^n; t^{n+1}, x), t^n)$

Method of Characteristics



Flux-Balance Methods

- Use characteristics for finite-volume:

$$\int_{x_{i-1/2}}^{x_{i+1/2}} f(t^{n+1}, x) dx = \int_{X(t^n; t^{n+1}, x_{i-1/2})}^{X(t^n, t^{n+1}, x_{i+1/2})} f(t^n, x) dx$$

- Define flux terms:

$$\Phi_{i+1/2}(t^n) = \int_{x_{i+1/2}-v\Delta t}^{x_{i+1/2}} f(t^n, x) dx$$

Flux-Balance Methods

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- Rewrite first equation as

$$\int_{x_{i-1/2}}^{x_{i+1/2}} f(t^{n+1}, x) dx = \int_{x_{i-1/2}}^{x_{i+1/2}} f(t^n, x) dx + \frac{\Phi_{i-1/2}(t^n) - \Phi_{i+1/2}(t^n)}{\Delta x}$$

Flux-Balance Scheme

- Leads to the finite-volume fully-discrete scheme

$$f_i^{n+1} = f_i^n + \frac{\Phi_{i-1/2}(t^n) - \Phi_{i+1/2}(t^n)}{\Delta x}$$

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- How to evolve the scheme? Need to compute fluxes Φ .
 - No time-step restriction!
 - Explicitly compute integral definition (reconstruct f)
 - Compute pointwise primitive values (reconstruct primitive, F)

The Semi-Lagrangian Method

- Very similar in spirit to flux-balance method: follow the characteristics
- Degrees of freedom: point-values $f^n(x_i)$
- Recall:

$$\frac{dX(s)}{ds} = v(s, X(s))$$

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$$f(x, t^{n+1}) = f(X(t^n; t^{n+1}, x), t^n)$$

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- The semi-Lagrangian method:

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Semi-Lagrangian Schemes

- We have $f^n(x_i)$ for all i

Semi-Lagrangian Schemes

- We have $f^n(x_i)$ for all i
- Need to interpolate to $X(t^n; t^{n+1}, x_i) = x_i - v\Delta t$ to advance the scheme
 - Lagrange interpolations
 - Implementable to high-order interpolants
 - Conservation: only with a central stencil, linear constant-coefficient advection PDE
 - Hermite interpolations
 - Match function values and cell interface derivative values
 - Conservation: same as Lagrange interpolants

Other Solvers

- Spectral methods – computational savings: FFT
- Finite-difference methods – Conservation: mass, energy

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- Spectral methods – computational savings: FFT
- Finite-difference methods – Conservation: mass, energy
- DG methods – positivity

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The Collision Operator–Steady-State

Evolves towards ~~Gaussian~~ *in v* *Maxwellian*

$$f_{\infty}(v) := \frac{\rho}{(2\pi T)^{d/2}} \exp\left(-\frac{|V - v|^2}{2T}\right)$$

- Steady-state: $Q(f_{\infty}, f_{\infty}) = 0$

Simplifying the Collision Operator

$$Q(f, f) = \int_{\mathbb{R}^3} \int_{S^2} B [f(v')f(v'_*) - f(v)f(v_*)] d\omega dv_*$$

Linearize: $f := f_\infty + g$

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$$Q^{\text{lin}}(f, f) \approx \int_{\mathbb{R}^3} \int_{S^2} B [g(v')f_\infty(v'_*) + f_\infty(v')g(v'_*) \\ - g(v)f_\infty(v_*) - f_\infty(v)g(v_*)] d\omega dv_*$$

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$$\begin{aligned} Q^{\text{lin}}(f, f) &\approx \int_{\mathbb{R}^3} \int_{S^2} B [g(v')f_\infty(v'_*) + f_\infty(v')g(v'_*) \\ &\quad - g(v)f_\infty(v_*) - f_\infty(v)g(v_*)] d\omega dv_* \\ &\rightsquigarrow \int_{\mathbb{R}^3} k(v, v') f_\infty(v) f(v') - k(v', v) f_\infty(v') f(v) dv' \end{aligned}$$

→ *Linear Vlasov-Boltzmann Equation*

Simplifying the Collision Operator

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Even simpler: *Bhatnagar-Gross-Krook (BGK) Approximation*

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$$Q^{\text{BGK}}(f) = \frac{1}{\tau} (f_\infty - f)$$

Discretizing Velocity Space

v -space: infinite. Computers: finite.

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$$\mathbf{1} \quad \text{supp}_v(Q(f, f)(v)) \subset \mathcal{B}(0, R\sqrt{2}).$$

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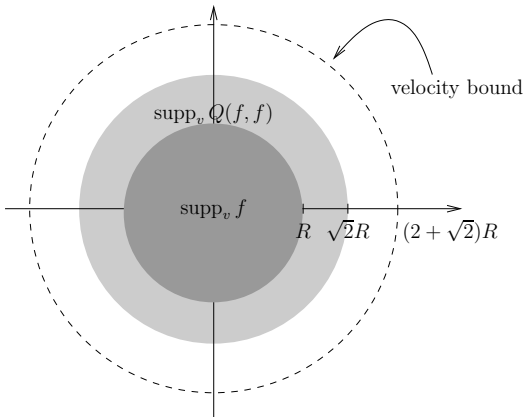
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- 1** $\text{supp}_v(Q(f, f)(v)) \subset \mathcal{B}(0, R\sqrt{2})$.
- 2** *Change of variables $g := v - v_*$. We can restrict $g \in \mathcal{B}(0, 2R)$ with no change:*

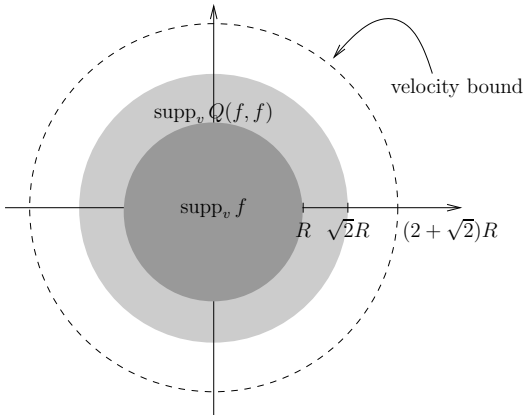
$$Q = \int_{\mathcal{B}(2R) \times S^2} B(|g|, \theta) [f'(v')f(v'_*) - f(v)f(v-g)] d\omega dg$$

Compact support for Q



$\text{supp}_v(f)$ grows by a factor of $\sqrt{2}$ each timestep

Compact support for Q



$\text{supp}_v(f)$ grows by a factor of $\sqrt{2}$ each timestep—for any $\Delta t > 0!$

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Hope: Similar savings!

→ Investigate Fourier Transforms in v -Space.

Fourier Transform in v -Space

Introduce a lattice:

$$v_k := h_v k, \quad h_v > 0, \quad k \in \mathbb{Z}^3$$

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using the midpoint rule

$$\tilde{\varphi}(\xi) := h_v^3 \sum_{k \in \mathbb{Z}^3} f(v_k) e^{i(v \cdot \xi)}.$$

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FFT requires: $h_v h_\xi = 2\pi/n$ for $n \in \mathbb{N}$.

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Get n -periodicity in k and j :

Fourier Transform in v -Space

$$\tilde{\varphi}(\xi) := h_v^3 \sum_{k \in \mathbb{Z}^3} f(v_k) e^{i(v \cdot \xi)}.$$

But for which ξ do we need to evaluate this?

Pick a grid in ξ -space, too: $\xi_j := h_\xi j$ with $h_\xi > 0$ and $j \in \mathbb{Z}^3$.

FFT requires: $h_v h_\xi = 2\pi/n$ for $n \in \mathbb{N}$.

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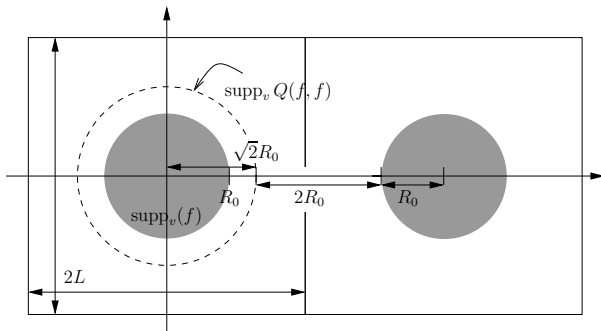
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Faking Periodicity

So v -space becomes (artificially) $2L$ -periodic, with $L = nh_v/2$.

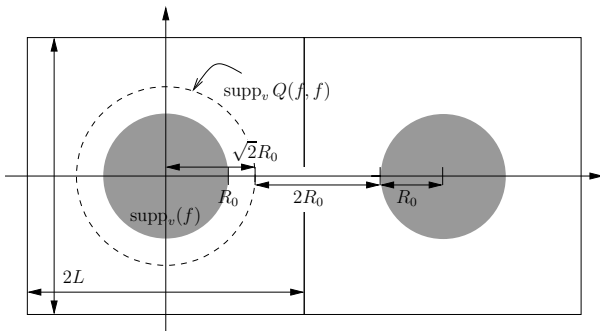
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Avoid aliasing: choose $L \geq \frac{1}{2}(\sqrt{2}R_0 + 2R_0 + R_0) = \frac{3+\sqrt{2}}{2}R_0$.

Overview

	M	H	V	O	Order	Ops	Ms	Mm	En
DSMC	✓	✓	✓	✓	$N^{-1/2}$	N^3	✓	✓	✓
IbrR	✓	✓	✓	—	N^{-2}	N^6	✓	○	○
gen.	—	—	—	✓	N^{-2}	$N^6 \log N$	✓	○	○
BobR	—	3	—	—	N^{-2}	$N^6 \log N$	✓	○	○
ParRu	✓	✓	✓	—	$N^{-\infty}$	N^6	✓	○	○
MouP	2	3	—	—	$N^{-\infty}$	$N^5 \log N$	✓	○	○

Bobylev-Rjasanow's Integral Transform Method

Idea:

Rewrite $Q(f, f)$ into convolution integrals.

Exploit the Fast Fourier Transform to evaluate these integrals quickly.

▶ skip the details

Bobilev-Rjasanow's Method

Lemma (Symmetric Representation)

The collision operator can be represented as

$$Q(f, f) = 4 \int_{\mathbb{R}^3} \int_{\mathbb{R}^3} \tilde{B}(y, z) \delta(z \cdot y) [f(v + z)(f(v + y) - f(v)f(v + y + z))] dy dz,$$

with

$$\tilde{B}(y, z) := \frac{1}{|y + z|} B \left(|y + z|, \frac{y \cdot (y + z)}{|y||y + z|} \right).$$

▶ skip the rest of this method

Bobilev-Rjasanow's Method

Theorem (Integral Transform Representation)

The collision operator for hard spheres can also be represented as

$$Q(f, f) = \int_{S^2} \int_{S^2} \delta(\omega_1 \cdot \omega_2) [\Phi(\omega_1)\Phi(\omega_2) - f\Psi(\omega_1, \omega_2)] d\omega_1 d\omega_2,$$

(note the implicit v -dependency)

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Remarks on Bobylev-Rjasanow

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Apply cleverness to exploit structure.

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- Ψ : "*Generalized Radon Transform*"
- Φ and Ψ of convolution type: good for FFT.

Rewriting X-Ray Transform \rightarrow Fourier Transform

$$\Phi(v, \omega) = \mathcal{F}^{-1}[\mathcal{F}[\Phi](\xi, \omega)](v, \omega)$$

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The Algorithm

$$Q(f, f) = \int_{S^2} \int_{S^2} \delta(\omega_1 \cdot \omega_2) [\Phi(\omega_1)\Phi(\omega_2) - f\Psi(\omega_1, \omega_2)] d\omega_1 d\omega_2$$

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Altogether: $N^6 \log N$

Pareschi-Russo's Spectral Method

Idea:

Do Fourier-Galerkin on $\partial_t f = Q(f, f)$.

Demand that the residual of $\partial_t f_N = Q(f_N, f_N)$ is orthogonal to all trigonometric polynomials, where f_N is a Fourier expansion.

▶ skip the details

Pareschi-Russo's Spectral Method

Expand

$$f_N(v) := \sum_{k \in K_N^3} \hat{f}_k e^{i(k \cdot v)}$$

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→ get the scheme

$$\frac{\partial \hat{f}_k}{\partial t} = \sum_{m \in K_N^3} \hat{f}_{k-m} \hat{f}_m \hat{B}(k-m, m) - \sum_{m \in K_N^3} \hat{f}_{k-m} \hat{f}_m \hat{B}(m, m)$$

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with the *kernel modes*

$$\hat{B}(l, m) = \int_{B(0, 2R_0)} \int_{S^2} B(|g|, \cos \theta) e^{-ig \cdot \frac{(l+m)}{2} - i|g|\omega \cdot \frac{(m-l)}{2}} d\omega dg$$

Performance of the Pareschi-Russo Spectral Scheme

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 - Recall: naive method would take $O(N^6 N_a)$.

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Performance of the Pareschi-Russo Spectral Scheme

$$\frac{\partial \hat{f}_k}{\partial t} = \sum_{m \in K_N^3} \hat{f}_{k-m} \hat{f}_m \hat{B}(k-m, m) - \sum_{m \in K_N^3} \hat{f}_{k-m} \hat{f}_m \hat{B}(m, m)$$

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 - Recall: naive method would take $O(N^6 N_a)$.
- Big plus: Rolled evaluation of problematic sphere integral into pre-computation.
- Loss term is sum of convolutions.
 - Can evaluate in $O(N^3 \log N)$.

Kernel modes of the Pareschi-Russo Spectral Scheme

$$\hat{B}(l, m) = \int_{\mathcal{B}(0, 2R_0)} \int_{S^2} B(|g|, \cos \theta) e^{-ig \cdot \frac{(l+m)}{2} - i|g|\omega \cdot \frac{(m-l)}{2}} d\omega dg$$

- The pre-evaluated \hat{B} takes up only limited storage.

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→ Two-dimensional array suffices.
- A far more explicit expression for \hat{B} is available.

Analysis of the Pareschi-Russo Spectral Scheme

Theorem (Spectral Accuracy)

Let $f \in L^2([-\pi, \pi]^3)$. Then

$$\|Q^{\text{tr}}(f, f) - \mathcal{P}_N Q^{\text{tr}}(f, f)\|_2 \leq C \left(\|f - f_N\|_2 + \frac{\|Q^{\text{tr}}(f_N, f_N)\|_{H_{\text{per}}^r}}{N^r} \right)$$

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Can show stability after including some filtering in the scheme.

Mouhot-Pareschi's Spectral Method

Idea:

Combine the symmetric representation from Bobylev-Rjasanow with the Fourier-Galerkin approach from Pareschi-Russo.

▶ skip the details

Symmetric Representation

Recall

Lemma (Symmetric Representation)

The collision operator can be represented as

$$Q(f, f) = \int_{\mathbb{R}^d} \int_{\mathbb{R}^d} \tilde{B}(y, z) \delta(z \cdot y) [f(v+z)f(v+y) - f(v)f(v+y+z)] dy dz,$$

with

$$\tilde{B}(y, z) := \frac{2^{d-1}}{|y+z|^{d-2}} B\left(|y+z|, \frac{y \cdot (y+z)}{|y||y+z|}\right).$$

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(slight generalization for multi- d)

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Fourier-Galerkin Derivation

▶ skip the details

$$0 \stackrel{!}{=} \int_{[-\pi, \pi]^d} \left(\frac{\partial f_N}{\partial t} - [Q^+(f_N, f_N) - L(f_N)f_N] \right) e^{-ik \cdot v} dv$$

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Remarks on the Scheme

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Pareschi-Russo scheme looked exactly the same.

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Recall: Pareschi-Russo loss term could be evaluated in $O(N^3 \log N)$ because of special structure.

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- Discretize

$$\theta := p\pi/M, \quad \varphi := q\pi/M.$$

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→ Sum of M^2 convolutions.

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$$\begin{aligned}
 \frac{\partial \hat{f}_k}{\partial t} &= \frac{\pi^2}{M^2} \sum_{l,m \in K_N^3} \hat{f}_{k-m} \hat{f}_m \beta(k-m, m) - \text{loss} \\
 &= \frac{\pi^2}{M^2} \sum_{m \in K_N^3} \hat{f}_{k-m} \hat{f}_m \sum_{p,q=0}^M \alpha_{p,q}(k-m) \alpha'_{p,q}(m) - \text{loss} \\
 &= \frac{\pi^2}{M^2} \sum_{p,q=0}^M \sum_{m \in K_N^3} \hat{f}_{k-m} \alpha_{p,q}(k-m) \hat{f}_m \alpha'_{p,q}(m) - \text{loss}.
 \end{aligned}$$

→ Sum of M^2 convolutions. Use FFT: $M^2 N^3 \log N$ operations.

Final Remarks on Mouhot-Pareschi

$$\frac{\partial \hat{f}_k}{\partial t} = \frac{\pi^2}{M^2} \sum_{p,q=0}^M \sum_{m \in K_N^3} \hat{f}_{k-m} \alpha_{p,q}(k-m) \hat{f}_m \alpha'_{p,q}(m) \quad - \quad \text{loss.}$$

- Fastest method around.

